# STUDY OF TEMPERATURE HETEROGENEITIES AT SUB-KILOMETRIC SCALES AND INFLUENCE ON SURFACE-ATMOSPHERE ENERGY INTERACTIONS

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Abstract— The retrieval of land surface temperature  $(LST)^2$ 2 from remote sensing techniques has been studied and validate $\hat{d}^3$ 3 4 during the past forty years, leading to important improvements<sup>34</sup> 5 Accurate LST values are currently obtained through measure<sup>35</sup> 6 ments using medium-resolution thermal infrared (TIR) sensors<sup>36</sup> 7 However, the most recent review reports demonstrated that 7 8 future TIR LST products need to obtain reliable temperature8 9 values at a high spatial resolution (100 m or higher) to study 10 temperature variations between different elements in a heterogeage 11 neous kilometric area. The launch of high-resolution TIR sensors 12 in the near future requires studies of the temporal evolution and 13 spatial heterogeneities of the elements in a mixed region. The 14 present study analyzes the LST in a sub-kilometric highly heter42 15 ogeneous area, combining the use of LST products from 3 high-resolution TIR orbiting sensors with the LST maps created  $\underline{A}$ 16 from a TIR camera onboard an unmanned aerial vehicle  $(UAV)_{45}^{++}$ 17 18 The aim is to estimate the LST variability in a heterogeneous area containing different surfaces (roads, buildings, grass, etc.),6 19 20 observed from different TIR sensors at different spatial resolu47 21 tions, covering from the meter to the kilometer scales. Several8 22 results showed that variations in the LST up to 18 °C were identiag fied with the UAV-TIR camera, and significant differences were 23 also present in the LST products obtained from simultaneous 1 overpasses of high-resolution satellite TIR sensors. A second 24 25 objective of the study, due to the availability of the high $5^2$ 26 27 resolution LST fields, was to explore the thermal advection be<sup>53</sup> 28 tween the different elements and determine if it correlates with 4 29 the surface energy budget in the same area, thus indicating that5 this process is of importance for heterogeneous terrains at these 30

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scales. This study also highlights the relevance of the UAV-TIR camera flight for future studies since it is not commonly used in TIR remote sensing but has substantial potential advantages.

*Index Terms*—Land surface temperature, Remote sensing, Surface heterogeneity, Surface-atmosphere exchange fluxes, Unmanned aerial vehicle, Thermal advection.

# I. INTRODUCTION

AND surface temperature (LST), recognized as one of the →most important Earth System Data Records according to NASA [1], requires accurate estimations with small uncertainty and bias [2]. LST is an important variable that controls processes in several disciplines, such as agro-meteorology [3] or surface-atmosphere interactions [4]. However, recent review reports have stated that in addition to the accurate retrievals of LST, the future of thermal infrared (TIR) remote sensing should use LST products at high spatial resolutions (above 100 m) to obtain reliable temperature values from the different elements composing a heterogeneous kilometric area, considering that the temperature of such elements is more important than the area average value [5]. In previous works [19], it has been proposed that the standard deviation of LST values over a heterogeneous area could be used as a parameter to show the variability in LST. In any case, the LST products from medium resolution TIR sensors (~  $1 \text{ km}^2$ ) are very reliable and accurate ( $\pm 1$  °C), as shown by the comparisons with sampling field data measured in highly homogeneous surfaces [6-7]. However, it has been demonstrated that such kilometric LST products are not representative of the temperature of different elements on heterogeneous surfaces [8, 65].

Recent studies indicate the necessity of obtaining LST maps from satellite TIR sensors at a high spatial resolution (60-100 m) to carry out work on heterogeneous zones, such as the monitoring of the urban heat island effect and control of maximum temperatures in cities [9-10], the classification of different surface types to improve irrigation practices [11], the determination of patterns in the temporal evolution of the LST

[12, 70], the management of strategies for forest recovery and a 69 70 reforestation after a fire based on the high sensitivity of the? 71 LST to different severity levels of the burned zones [13], the 28 estimation of evapotranspiration (ET) maps at spatial resolu29 72 73 tions lower than 100 m [14] or the study of the impact of theke0 74 LST high-resolution maps on the surface fluxes [15]. 131 75 LST maps at fine scales (below 10 m) have been recentl \$2 76 used in some studies by means of TIR cameras mounted 183 77 airborne systems, such as helicopters or airplanes, to performate 78 detailed studies of LST heterogeneities associated witB5 small-scale components present in a hectometer area [16B6 79 80 These studies corroborated that LST products at such fines?

scales could be useful to retrieve accurate LST values of the
leaf in a crop [16, 71], which allows good episodic or seasonka
estimations of the ET for this crop [14].

84 Unmanned aerial vehicles (UAV) are able to determine LST 1 85 and atmospheric variables. They can make several dail\$2 86 flights. The actual surface that is sampled depends on the 43 vehicle and the legislation that usually limits operation 144 87 88 visual control of the drone by the ground-based operator tol45 89 radius of a few hundred meters. The speed of the UAV 1i46 90 usually lower than manned aircraft and allows a higher spatial7 91 resolution, discriminating the variations in LSTs at very find 8 92 scales. The use of UAVs in remote sensing applications hh49 increased significantly in the last triennium [17]. These appli50 93 94 cations include a variety of disciplines, such as agriculture51 95 forestry or geology (an interesting list of published studies caff2 96 be consulted in Table I of [18]). 153

97 However, there are very few works about the use of TIE4 98 sensors installed in a UAV, and there is room to explore 1f5 99 depth the possibilities of such technology in the TIR remote6 100 sensing field. Some of these studies were motivated, for ih57 101 stance, to estimate the importance of the thermal advection kat8 different scales [19], to evaluate the thermal impact of a rives9 102 103 on the surrounding agricultural terrain [20] or to compar60 104 turbulent energy flux estimates with two surface energy bal61 105 ance (SEB) models based on UAV TIR imagery with high 106 spatial and temporal resolutions [21]. Other types of applicate tions include the accurate mapping of the geothermal heat 107 108 signature of a geothermal field [22] and monitoring animal populations for conservation purposes [23]. A review of for-109 estry applications of UAVs in Europe with thermal camerals 110 166 111 can also be found in the literature [24].

Having access to the variability of LSTs allows the estima  $\frac{167}{100}$ 112 tion of thermal advection between distinct surface elements of 113 the study area. The surface energy budget considers all the 169114 energy exchanges that take place in the atmosphere-surface 115 interface. In this framework, the four main terms (i.e., the net 116 Ž2 radiation (Rn), turbulent sensible (H) and latent (LE) heat 117 fluxes and the soil (G) heat flux) would be balanced if they are  $\frac{173}{12}$ 118 observed over a homogeneous surface (and thus with no ther-119 mal advection) and if there are no other thermal sources  $b_{1}^{75}$ 120 sinks. In contrast, with LST heterogeneities, the thermal  $a_{-}^{176}$ 121 vection may become important, and the traditional 4-term 122 energy budget would not close. It has been [25] shown that 123 using two years of data over heterogeneous surfaces, the im<sup>79</sup> 124 balance of the surface energy budget was usually between  $10^{10}$ 125

to 30% of the  $R_n$ , often being the second or third term in importance and overpassing the ground or turbulent heat fluxes, especially at night. These values are in accordance with those provided in the review of [26]. The authors indicated that there were a number of sources to explain the imbalance, including instrumental disposition, lack of stationarity, or missing processes such as biological or anthropic activity or thermal advection. A first estimation of the importance of the advection term was given in [19] using model, satellite and UAV TIR data, where they showed that hectometer scales were likely to contribute significantly to the SEB imbalance if there were persistent heterogeneities present.

A recent study conducted in the Campus of the University of the Balearic Islands (UIB) has shown that the different elements composing this area form a temperature field with large heterogeneities [27]. In that work, the MODIS values for the pixel covering the campus are compared to ground measurements and indicate a larger bias and RMSE values than for similar studies in homogeneous terrain. Furthermore, an analysis of the Landsat and ASTER images indicate large surface variability in LSTs in the range of 10 to 20 °C depending on the season. Such an analysis has been key to plan the current work and a companion study by [28]. In the current case, the main objective is to study the surface temperatures at subkilometric scales through a combined inspection of satellitederived LST products and observations from an UAV. In addition, these fields are used to explore the contribution of such heterogeneities in the imbalance of the SEB.

The paper proceeds as follows: Section II describes the heterogeneous site, the materials and the methods selected for the study. In Section III, the LST products observed from different platforms are presented and validated, while the corresponding heterogeneity results are discussed. An application study estimating the importance of the thermal advection with the SEB equation is demonstrated in Section IV. Finally, conclusions are provided in Section V.

# II. STUDY SITE AND MATERIALS

#### A. UIB Campus

The study site is located in the campus of the University of the Balearic Islands (UIB), at a height of 80 m above sea level (a.s.l.) in the Palma Basin at Mallorca Island (Figure 1), in the western Mediterranean Sea and 200 km east from the Iberian Peninsula. The UIB campus has an approximate area of 1 km x 1 km, halfway between the city of Palma and the natural area of the Tramuntana Mountain range. It is composed of many different types of surfaces such as buildings, asphalted roads, farming areas, some sloping terrain, fields of orange and almond trees, lawns and some natural vegetation extensions. The natural vegetation includes wild grass between October and May that dies in the summertime, with the surface becoming a mixture of bare soil and dead vegetation.

Recent studies showed the heterogeneity of the UIB Campus in terms of LST with gradients between different surfaces of up to 10 °C in wintertime and up to 20 °C in summertime, during the daytime [27]. During weak large-scale pressure gradients and clear-sky conditions, locally generated winds area
present in Mallorca and especially in its three main basins [2935
30]. This scenario is the case for the diurnal sea breeze (especial
cially from April to October) or the nocturnal land breeze;
both often coupled with slope winds [31].

186 A complete surface energy budget (SEB) station has bears 187 running at the UIB Campus since 2015 (yellow dot in Figure 40 188 1). Several data taken for the entire year of 2016 from the taken the taken for the entire year of 2016 from the taken the taken the taken the taken taken the taken taken the taken 189 station are used in the present study. First, observations from 242 190 broadband thermal infrared radiometer (IR120, Campb2H3 191 Scientific) were taken as a ground reference value to validate4 192 several LST products (see Section III). Second, the measur245 193 terms of the SEB together with the wind observations are used6 194 in Section IV. These contributions to the SEB are the turbule247 sensible and latent heat fluxes, the net radiation and t248 195 ground heat flux. Turbulent fluxes were measured with a sonic 249196 197 anemometer (81000, RM Young) and a gas analyzer (EC150 Campbell Scientific) after applying the Eddy correlation tech  $\frac{250}{100}$ 198 nique [32]. The net radiation was extracted from the upward<sup>1</sup> 199 and downward components of the solar and longwave radra52 200 tion observed from the Hukseflux NR01 net radiometer. Finat 201 ly, the ground heat flux was determined as the contribution  $\frac{25}{6}$ 202 the measured flux by the Hukseflux HFP01 sensor at a certain<sup>5</sup> 203 depth within the soil plus the change in heat storage in the  $s^{2}$ 204 above the plate [33], which was estimated using soil tempera<sup>257</sup> 205 ture and water content measurements [34]. The IR120 radiom<sup>58</sup> 206 eter operates in the 8-14  $\mu m$  spectral range and has an FOV  $\frac{259}{25}$ 207 36°. This sensor is located at a 1-m height, measuring the  $LSP^{0}$ 208 261 209 representative of a circumference with 2 m of diameter.

210 In situ LST (in °C) is calculated from the measured 2 211 upwelling longwave radiance ( $L^{\uparrow}$ , in Wm<sup>-2</sup>) as: 263 264

212 
$$LST = \left[\frac{L^{\uparrow} - (1 - \varepsilon)L^{\downarrow}}{\varepsilon\sigma}\right]^{1/4}$$
(1)266  
267

213 where  $\varepsilon = 0.97$  is the selected value for the broadband emis68 214 sivity surface [35], corresponding to senescent sparse shrub69 215 and  $\sigma = 5.67 \cdot 10^{-8} \text{ Wm}^{-2}\text{K}^{-4}$  represents the Stefan–Boltzmain0 216 constant. Reflected downwelling radiance (L<sup>1</sup>, also in Wm27)1 217 is calculated from the Brutsaert approximation [36], as it 232 218 done in [28]. 273

219 Ground-based LST observations correspond to the 1-man4 220 average measurements. The IR-120 radiometer was calibrated5 in the laboratory against the reference blackbody model 221 LANDCAL P80P. In 2009, the P80P blackbody participated 222 in a comparison campaign organized by the Committee  $\delta n'$ 223 Earth Observation Satellites in the National Physical Labora-224 tory (NPL). The results showed that the P80P blackbody  $^{2}$ 225 226 agreed with the NPL reference radiometer within  $\pm 0.15$  °C a  $^{20}$  20 -30 °C [37]. The reference blackbody temperature range 227 within 0-60 °C, showing an RMSE on the retrieved LST from 283 228 229 the IR -120 radiometer data of  $\pm 0.4$  °C. 284

# 230 B. MODIS data

231 The Moderate Resolution Imaging Spectroradiometers
232 (MODIS) onboard the Terra and Aqua satellites current
233 provides global coverage, twice-daily LST, and emissivities

generated from different algorithms: products two MOD11 L2 (MYD11\_L2 for Aqua) MOD11B1 and (MYD11B1 for Aqua). The MOD11\_L2 product uses the generalized split-window algorithm [38] and provides LSTs and classification-based emissivities for bands 31 (10.8-11.3 μm) and 32 (11.8–12.3 μm) at 1-km resolution. In this study, the LST product used was the MOD11 L2 with refinements proposed in its version 6 [39]. Two images (day and night) of the MOD11\_L2 LST product were downloaded from the NASA webpage (https://reverb.echo.nasa.gov) for July 21st, 2016. These scenes were acquired in cloudless conditions, and the associated LST uncertainty included in the product stipulates a value of 1 °C. This uncertainty is only an estimated value, but it is quite conservative in real clear-sky conditions [39].

### C. ASTER data

The Advanced Spaceborne Thermal Emission and Reflection Radiometer (ASTER, [40]) is also a MODIS onboard the sun-synchronous Terra satellite. The ASTER uses five TIR bands to measure the Earth's emittance within the 8-13 µm range, offering an LST and emissivity ( $\epsilon$ ) product at 90 m x 90 m spatial resolutions every 16 days but with an 8% duty cycle, after applying the semi-empirical temperature and emissivity separation (TES, [41]) method. The uncertainties associated with LST and  $\varepsilon$  after the TES method are 1.5 °C and 0.015, respectively. The retrieval of the ASTER LST from the TES shows inaccurate estimates over surfaces with low emissivity spectral contrast or under humid atmospheric conditions [42]. To minimize the LST errors associated with atmospheric correction, [43] proposed a water vapor scaling (WVS) method that improves the accuracy of the water vapor atmospheric profiles on a band-by-band basis for each observation. Implementing both the WVS and TES methods showed a significant improvement in the retrieved LST [44, 6].

ASTER-Terra overpasses the UIB Campus twice a day, every 16 days, approximately 1049 UTC (morning scene) and 2153 UTC (night scene). A total of 11 (6 at morning, 5 at night) cloudless ASTER images were acquired between April and November 2016 over the UIB Campus. Table I shows the date and UTC hour of the cloudless ASTER scenes used in this study. ASTER image acquisition is an on-demand service, so we requested collection images over our Mallorca site.

#### D. Landsat 7- ETM+ data

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The Enhanced Thematic Mapper Plus (ETM+) onboard the heliosynchronous Landsat 7 covers the entire Earth in 16 days. The ETM+ measures the radiance in eight spectral bands ranging from the visible spectrum to the TIR range with a spatial resolution of 30 m x 30 m. The TIR band 6 (10-12  $\mu$ m) is disaggregated from 60 m to 30 m by the Landsat team [45]. A failure in the scan line corrector (SLC-off mode) occurred in 2003, and since then, the ETM+ images have been affected by a 100-m strip line of void-data, every kilometer. The location of Mallorca Island allows an increase of the overpass frequency of the Landsat 7-ETM+ to every 7–9 days, at approximately 1030 UTC because the island is placed between the pass of

two different orbits of Landsat 7-ETM+, which scan the islafield
from the north (path/row 197/32) and the south (path/row 196/33).
Figure 2 shows an RGB view of both Landsat 3746
ETM+ overpasses over Mallorca Island.

293 The retrieval of the LST from Landsat 7-ETM+ is based **94**7 294 the single-channel method [46], which applies atmospheric 295 and surface corrections to the top of atmosphere (TOA) spec-296 tral ETM+ radiance ( $L_{TOA,i}$ , in Wsr<sup>-1</sup>m<sup>-2</sup>µm<sup>-1</sup>) measurement 297 at band 6. The LST is cleared from the radiative transfer equat 298 tion (RTE) in (2): 350 351

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$$L_{TOA,i} = \left[\varepsilon_i B_i(\text{LST}) - (1 - \varepsilon_i) L_{hem,i}^{\downarrow}\right] \tau_i + L_{atm,i}^{\uparrow}$$
 (2)352

353 where  $\varepsilon_i$  is the surface emissivity;  $B_i(LST)$  is the Planes  $\overline{\xi_4}$ 300 function of a blackbody emitting at the surface temperature 355301 (LST); and  $L^{\downarrow}_{hem,i},\,\tau_i$  and  $L^{\uparrow}_{atm,i}$  are the atmospheric parameters  $_{356}$ 302 corresponding to hemispherical downwelling radiance, atm557303 phere transmissivity and upwelling radiance, respectively. The 304 subscript i refers to the channel-effective quantity of  $eaging_{9}$ 305 306 parameter in the RTE. 360

The surface emissivity used in Equation (2) was extracted 361307 from the ASTER Global Emissivity Database (GED) [47]. The 352308 ASTER-GED offers surface emissivity values at a 100-363 309 spatial resolution for the five TIR channels of the AST $\tilde{E}_{364}^{\tilde{R}}$ 310 sensor based on planetary ASTER TES results data from  $20\tilde{9}_{55}$ 311 to 2008. In this study, the emissivity used to correct the sure  $\frac{3}{366}$ 312 face emission at the disaggregated TIR image of the ETM<sub>467</sub> 313 314 sensor at 30 m  $\times$  30 m pixels was calculated from the8 315 weighted averaging ASTER GED values in channels 89  $(10.25-10.95 \,\mu\text{m})$  and 14 (10.95-11.65  $\mu\text{m}$ ) to the ETM+ TLR  $_{370}$ 316 317 band 6, since both channels cover the spectral resolution **3**71 such a band. After the convolution process, the 100 m x  $100 \frac{m}{372}$ 318 resolution data were downscaled to 30 m x 30 m. The emissi $\frac{1}{373}$ 319 320 ity of different components of the UIB Campus ranges be74 tween 0.960–0.982, with an average value of 0.972 and  $\frac{34}{375}$ 321 deviation of 0.004. According to [41], the associated emissivity  $\frac{1}{16}$ 322 ty uncertainty is approximately 0.015, and therefore, the  $LST_{77}$ 323 324 values might also have an uncertainty of approximately 1378 325 °C.

The atmospheric variables  $L^{\downarrow}_{hem,i}$ ,  $\tau_i$  and  $L^{\uparrow}_{atm,i}$  in (2) weight 326 calculated with the MODerate resolution atmospheric 327 TRANsmission (MODTRAN) radiative transfer code 328 3<sup>8</sup>2 5.2.1, [48]) using the synthetic atmospheric profile provid $\xi d_{33}^{-1}$ 329 by a web-tool calculator [49] based on the National Centers 330 for Environmental Prediction (NCEP) model [50]. Compared 331 with sounding data, this NCEP profile was demonstrated to  $\frac{1}{586}$ 332 the best option to retrieve atmospheric variables, with respect  $\frac{1}{387}$ 333 to other synthetic atmospheric profiles [51]. To retrieve $\frac{1}{388}$ 334 more accurate synthetic atmospheric profile, the web-tool 335 calculator allows establishing surface conditions of the select $_{50}$ 336 ed location. These surface parameters consist of the altitude 337 (km), pressure (mb), air temperature (°C) and relative humid $\frac{1}{392}$ 338 ty (%), and they were provided in this study at the correspondence  $f_{93}$ 339 ing UTC time by the SEB station (see section II.A and Figure 4 340 341 1). 395

Once the variable  $B_i(LST)$  was estimated in the RTE from 343 the surface emissivity and atmospheric effects, the LST could

be obtained following the expression proposed by [45], with a noise equivalent change in temperature (NE $\Delta$ T) at 280 K of 0.22 K, as:

$$LST = \frac{1282.71}{\ln(666.09/B_i + 1)}$$
(3)

A total of 19 Landsat 7-ETM+ clear-sky scenes (8 for the 196/33 and 11 for the 197/32 path/row overpass) of the UIB Campus site were used in this study along the year 2016 (see Table II).

# E. UAV-TIR camera data

A TIR camera was assembled in a UAV (Figure 3a) to reproduce an LST map of the UIB Campus, at high spatial resolution (2 m x 2 m) flying at a height of 200 m. The TIR camera model is the FLIR LEPTON® Long Wave Infrared [52]. Its dimensions are 11 x 12 x 6 mm and the FOV is 51° and 63.5° in the horizontal and diagonal views, which produces a TIR scene at different spatial resolutions depending on the height of the fly, due to the 80 (horizontal) x 60 (vertical) active pixels in each camera shot, respectively. This TIR camera measures the 6-15  $\mu$ m integrated value of the spectral radiance of a target. Figure 3b shows the spectral response of the FLIR LEPTON TIR camera between 6-15  $\mu$ m.

The TIR scene acquisition process consisted of flying the UAV-TIR camera system at a height of 200 m from the red cross in Figure 1 to the north (900 m distance), then turning west (70 m distance), from north to south (900 m distance), and then repeating the entire south-north-west-south process several times to cover the entire campus. The overall process took 20 minutes, including a mid-flight stop to change the battery. The TIR camera took an average of 750 scenes of the whole UIB Campus (a scene every flying second). Table III shows the different UAV flights carried out in the five intensive operative periods (IOPs) programmed in June and July of 2016. It was initially proposed to fly the UAV-TIR camera every 2 hours each IOP day, starting at 0400 UTC and finishing at 2000 UTC. However, due to the battery or UAV system problems, the plan could not be fulfilled completely.

An LST map of each TIR scene was obtained after converting the digital numbers (DN) measured by the FLIR camera into radiance (LFLIR) according to the manufacturer's indications. The LFLIR is a composition of different radiance terms, as in Equation (2). The atmospheric terms were obtained from introducing the NCEP synthetic profile [49], limited between geopotential heights of 0.08 km (altitude of UIB campus) and 0.28 km (height a.s.l. of UAV flight), into the MODTRAN code. The broadband emissivity was considered as a unique constant value of  $0.964 \pm 0.015$ , for all pixels of the UIB Campus. It must be taken into account that such emissivity is different from the 0.972 selected for the Landsat 7 ETM+ sensor (see section II.D) because the operational spectral range of both sensors are different. The five different  $\varepsilon_i$  (j=ASTER channels 10-14) values of the ASTER GED, up-scaled at 1-km<sup>2</sup> resolution [47], were used to retrieve a constant value according to the broadband emissivity expression proposed #\$0
[53]: 451

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(4454

400  $0.333\varepsilon_{13} + 0.146\varepsilon_{14}$ 

401 Average emissivity values and the corresponding error a456 402 in good agreement with the emissivity range observed in the 7 403 UIB Campus for the Landsat 7- ETM+ data (see subsection58 II.D). No directional effects on emissivity [66-68] were co459 404 sidered in this study. Once every single TIR scene was co460 405 406 verted to the LST map, they were composed as an overall LST 1 407 map of the UIB Campus with the privative software Agiset 2 408 PhotoScan®. 463

The FLIR LEPTON TIR camera was also calibrated in t464 409 410 laboratory against the reference LANDCAL P80P blackbod §5 411 (see section II.A), showing an RMSE on the retrieved TH66 412 camera data of  $\pm 2$  °C. According to [49], associated errors 467 413 synthetic atmospheric profiles are  $\pm 2$  °C in air temperature a468  $\pm 2$  % in relative humidity. These uncertainties applied to t469414 atmospheric profiles used in this study and induced a corre- $\frac{1}{470}$ 415 sponding average uncertainty in the atmospheric parameters 416  $L^{\downarrow}_{hem,i}, L^{\uparrow}_{atm,i}$  and  $\tau_i$  of  $\pm 0.06 \text{ Wm}^{-2} \text{sr}^{-1} \mu \text{m}^{-1}, \pm 0.09 \text{ Wm}^{-2} \text{sr}^{-1} \mu \text{m}^{-1}$ 417 and  $\pm 0.005$ , respectively. These uncertainties, together with<sup>2</sup> 418 the emissivity error of  $\pm 0.015$  and TIR camera temperature 419 calibration uncertainty previously mentioned, established 474 420 total uncertainty on the LST retrieved from the UAV-TIR<sup>5</sup> 421 camera of  $\pm 3$  °C. However, the uncontrolled environmental<sup>6</sup> 422 conditions of the camera sensor can consequently lead to  $un^{477}$ 423 certainties in the camera performance and in the final  $LST^8$ 424 479 425 retrieval. 480

# 426 III. LST HETEROGENEITIES AT SUB-KILOMETRIC SCALE 481 482

427 A. Local validation of LST scenes

428 The vegetated surface targeted by the IR-120 radiometer **48**429 representative of the surface surrounding the SEB station, **a48**430 it is even representative of the 42% of the surface of the U**48**431 Campus. Therefore, we considered that this LST could be us **48**432 as a reference to validate the LST product calculated by **t48**433 ASTER-Terra, Landsat 7-ETM+ and UAV-TIR Camera se **48**434 sors. 490

435 Figure 4 shows the validation results for the LST product 491 ASTER, ETM+ and TIR Camera compared with the IR-1292 436 437 radiometer of the UIB budget energy station for all the data3 438 obtained for the corresponding sensor in Tables I, III and I494 A comparison between the MODIS LST product and that 95 439 440 measured by the ground radiometer was not considered b496 441 cause the spotted surface measured with the IR-120 is n4997 442 comparable with the kilometric LST value derived by MODIS98 443 especially in a heterogeneous terrain [27]. The ASTER, ETM99 444 and TIR Camera LST product shows a RMSD with respect 500 445 the radiometer temperature of  $\pm 1.3$  °C,  $\pm 1.8$  °C and  $\pm 3.1$  °C(1 respectively, and a bias of -0.5 °C, -0.5 °C and -0.6 °C, respectively 446 447 tively. In the case of the ASTER sensor, an RMSD of  $\pm 1.4$  303 for both morning and night overpasses was observed. For  $t\mathbf{\overline{b}}\Theta 4$ 448 ETM+ LST product, RMSDs of ±1.6 °C and ±2.3 °C wef05 449

observed for the 197/32 and 196/33 path/row orbits, respectively.

The validation results of this study are consistent with previous validation studies. Thus, for the ASTER sensor, LST uncertainties of ±1-2 °C were observed [54-56]. In the case of the Landsat 7-ETM+, the validation results of this study were also in good agreement with a previous UIB Campus study [27] and with other past published works [57-58] where an RMSD of ±1-2 °C was found in the LST product derived from the band 6 of the ETM+ sensor. The RMSD of the UAV-TIR Camera LST product was the expected uncertainty according to the sensitivity study described in section II.E. However, it is worth noting that the TIR camera validation was only accurate for the grass surface (case of SEB station) since in this case, it was well-known that the soil emissivity LST maps of Figure 5 are severely affected by different surface emissivity values. Other sources of error influencing the final uncertainty of these LST products could be the differences in the type of TIR sensor or the heterogeneity of the surfaces present in the UIB Campus.

### B. Daily LST heterogeneity of the UIB campus

Once the retrieval methods of the LST product for the 4 different sensors considered (MODIS, ASTER, ETM+ and TIR Camera) have been described and the corresponding uncertainties of such products stipulated, the LST heterogeneity of the UIB Campus was assessed. As commented previously, some of the surface elements of the UIB Campus are assigned an inaccurate emissivity value of 0.964 (see section II.E). For instance, the rooves of the buildings are made with different kind of tiles, which can have emissivity values ranging between 0.93-0.96, or asphalt from roads, which can be assigned an emissivity value of 0.93. This scenario means that some of the LST values obtained from the TIR camera can be inaccurate by several Celsius degrees, and their true values are most likely higher than those estimated in such LST maps. Nevertheless, considering an emissivity value of 0.96, instead of 0.93, may underestimate the LST by approximately 2-3 K, much less than the LST difference considered in this study. Notably, these types of surfaces have a representative extension in the southern region of the UIB Campus, but they do not have an important impact on the advection study carried out in the SEB station since at high resolution within 10-80 m, the surrounding surface is grass and trees, and those have emissivity values close to the selected 0.964 value. In addition, for the lower resolution 90-200 m, the effective emissivity of such buildings is close to the selected emissivity value, according to data extracted from the ASTER GED [47].

However, to avoid an underestimation of the LST value due to an incorrect selection of the emissivity, it is observed that for the four cloudless IOPs and for the central hours of the day (0800-1800 UTC), the hottest point is located at the artificial grass of the soccer field, and the coolest point is usually located in a humid creek area, situated north of the UIB Campus. Both sites have an extension area greater or at the same scale order of 100 x 100 m<sup>2</sup>, and according to the emissivity data from the ASTER 100 m GED database and following

Equation (4), the broadband emissivities of the soccer field2 506 507 and the humid creek area are 0.966 and 0.968, respective 563 508 Those emissivities are close to the emissivity selected in set 64 tion II.E for the whole UIB Campus, so we focused the LST65 509 510 heterogeneity study during the central hours of the day 566 511 those two sites for reliable conclusions. However, there couffed7 be a cavity effect for the artificial grass of the soccer fiebb8 512 which would cause the humid creek to appear colder than the9 513 514 soccer field. 570

515 Figure 5 shows the LST maps of the UIB Campus for foor1 of the nine UAV-TIR camera flights (0400, 1000, 1200 affd2 516 517 2200, UTC) during IOP5 (21/07/2016). The registration prof573 518 lems (maps a, b, c and d do not stack) in these LST maps area 519 due to the limitation of the privative software when asset 575 520 bling the TIR images taken by the UAV-TIR camera. The LST6 521 variability shown in maps of Figure 5 is representative of the7 522 four IOPs carried out in cloudless conditions (IOP 1, 2, 3 and 523 5). These IOPs showed significant LST differences within the 9 524 campus from the first flight. The maximal LST variatio580 525 observed in each IOP ranged within 3-28 °C during the days1 526 They ranged 3-10 °C during the two first flights before the table 2 527 sunrise (0400-0600 UTC). The LST differences between the 528 soccer field (warm point) and humid creek area (cold point84 529 reached significant values of 6-28 °C during the central hours 5 of the day (0800-1600 UTC). The last three flights before the 530 531 sunset (1800-2200 UTC) showed a slightly decreased LSTB7 532 variability of 6-12 °C. The LST difference at the early hours 588 533 the morning (0400-0600 UTC) and later hours close to sunses9 534 (1800-2200 UTC) are usually attributed to the difference be90 535 tween the LSTs of roads (hot point) and rooves of some buil591 536 ings (cold point). It is worth mentioning again that these d592 537 ferences could be reduced by several Celsius degrees since 593 538 inaccurate emissivity value was assigned to these types 5994 539 surfaces. 595

540 The IOP 4 (14/07/2016) and the early hours of the IOP596 541 were the only cases with an overcast sky and showed that the 7 542 LST variability in the majority of the flights that day remain 508 543 stable, ranging within 3-5 °C (see later Figure 8). Therefore99 544 under the same radiant lighting conditions, the LST variability 0 545 of the heterogeneous surfaces in the UIB Campus seemed 601 546 be less significant but still needs to be taken into account. The2 547 IOP 4 showed the important advantage of using the UA603 since under cloudy conditions, there was no possibility 604 548 549 retrieving the LST products from the TIR satellite sensors. 605 606

# 550 C. LST heterogeneity from orbiting sensors

The simultaneous overpass of the three TIR sensors me $60^7$ tioned (MODIS, ASTER and ETM+) matched with a UA  $60^{08}$ TIR camera flight in two occasions during the studied periof approximately 1030 UTC of the 05/07/2016 (IOP3) affold 21/07/2016 (IOP5). 611

Figure 6 shows the LST product of the four sensors at their different spatial resolution for the IOP 5 (the shown area is their same shown in Figure 5, so the coordinates are removed). Their high LST heterogeneity observed by the UAV-TIR cameria product at a resolution of 2 m x 2 m (Figure 6d) was consideria ably reduced as seen by the spatial resolution of 30 m x 30 fmin

of the LST product calculated for the Landsat 7-ETM+ TIR data (Figure 6c), where almost all the cold pixels of the UAV LST map were effectively converted to temperate LST values as a consequence of the spatial degradation. The loss of information of the LST heterogeneities was more evident with the LST product offered by the ASTER sensor (Figure 6b) at 90 m x 90 m, and this heterogeneity was imperceptible with the MOD11 LST product (Figure 6a) at 1 km x 1 km.

Despite the loss of the LST information, with the high-resolution LST product of the ETM+ and ASTER sensors with regard to that offered by the UAV-TIR Camera, both sensors still showed significant LST variability. For instance, the maximum LST difference observed between the soccer field and the wet creek in the north of the UIB Campus is 16 °C for the UAV-TIR Camera product; however, the difference was reduced to 11 °C with the ETM+ and to 9 °C with the ASTER LST products, and both differences are still very significant.

Figure 7 shows the LST product offered by the night overpass of the sensors ASTER and MODIS, onboard the Terra platform, compared with the LST product calculated from the simultaneous UAV-TIR camera flight over the UIB Campus. Under those stably stratified thermal conditions at night, significant LST differences were seen, such as the 14 °C value between the artificial grass of the soccer field (cold point) and the road in the southwestern part of the UIB Campus (warm point). This difference was reduced to 7 °C with the LST product offered by the ASTER sensor, which is still very significant. The MODIS showed LST differences of 2 °C for the four pixels covering partially the UIB campus.

Figures 6 and 7 are a clear example that a high spatial resolution on LST products is of key importance to study heterogeneities in zones composed of different surface types. In both cases, it was observed that the pixels composed by dense vegetation, such as the wet creek, showed the coldest zones, and other pixels composed by tarmac roads, bare soil or the soccer field, made with artificial grass, showed the hottest points of the LST map at midday. At midnight, the presence of grass (even the artificial one of the soccer field) and the tiled roof of some buildings were the coldest points, and the influence of tarmac roads produced the hottest pixels of the map. These temperature differences in both cases remained in the LST product of the ETM+ and ASTER sensors, but the MODIS did not show them, since they corresponded to subkilometric spatial resolutions.

# *D.* Analysis of the variability of the LST fields over the UIB Campus

To further explore the temporal and spatial variability of the LST fields estimated from the UAV-TIR Camera, ETM+ and ASTER, the corresponding probability density functions (PDFs) were computed over the UIB Campus for all the studied IOPs. As it was described in [59], the shape of the PDF of any magnitude depends on the variability of its values that, together with the statistics computed from it, allows estimating the heterogeneity of the field [60]. Figure 8 shows the PDFs computed from the LST fields from the UAV-TIR Camera,

618 ETM+ and ASTER over the UIB-Campus for all IOPs at d675 619 ferent instants to cover the diurnal cycle. The temporal evol676 620 tion of the statistics is shown in Figure 9 to check if the vari677 621 bility of the LST fields is similar for all IOPs through t678 622 inspection of the standard deviation ( $\sigma$ , gives informati679

about the spread of the LST values within the UIB Campus)

and the skewness (S, points out the prevalence of values larges0
or smaller than the mean one, corresponding to negative a681
positive S values, respectively).

Figures 8 and 9 indicate that the analysis of the PDFs is 83 627 useful tool to compare the spatial and temporal variability esta 628 the different sources of the LST (as those shown in Figures 65 629 and 7 as an example of a day and night instant, respectively 630 631 used in this work for the studied IOPs. It was found that the diurnal cycle of the variability of the LST fields reported by 632 the UAV-TIR camera was similar for all IOPs (similar shapes) 633 in the PDFs in Figure 8 and very close values of  $\sigma$  and S  $\dot{a}$   $g_0$ 634 Figure 9). However, IOP 4 presented a distinct behavior with 635 statistics close to normal values (small  $\sigma$  and S close to 0). The  $\sigma_2$ 636 637 presence of clouds during this IOP reduced the variability of 638 the LST field and was probably linked to a decrease in the net 3 639 radiation at the surface that did not allow the growth of the 640 thermal heterogeneities due to the different surface properties 94 641 A similar pattern was also found for IOP 1 at 0400 UTC (Fig95 ure 8a), when the PDF was narrower than the others due to the 96 642 presence of clouds during the late night and early morning that 643 disappeared as the day advanced, resulting in similar PDEs 644 shapes as IOPs 2, 3 and 5 (for instance, at 1200 UTC, Figure 9 645 646 8c). 700

647 During the central hours of the day for clear-sky IOPs, the 648 standard deviation reached the largest values of the diurnal? 649 cycle (Figure 9a) in agreement with the results of anothers study carried out in the UIB Campus [27]. In addition, the4 650 651 PDFs were biased with negative skewness demonstrating that5 the most frequent values are those LST values were warmeno 652 than that of the mean. As a result, the analysis of all IOP67 653 showed that heterogeneities were largest during the day, and 654 they were linked to the different surface properties of they 655 656 ground within the UIB Campus (Figure 5). During the morph 657 ing and evening transitions,  $\sigma(LST)$  took the minimal values of the day (minimum in the temporal evolution of  $\sigma$ , Figure 2 658 9a) but with a predominance of points with the LST colder3 659 660 than that of the mean (positive S, Figure 9b), demonstrati $\pi g_4$ that most of the surface covers at the UIB Campus had a fast5 661 662 response when the net radiation was close to zero. Finally 6 during the nighttime, the variability of the LST was larger than 7 663 664 during the transitions but was lower than during the day with 8 665 the equilibrium of the colder and warmer points (S close 709 666 zero). 720

667 The analysis of the PDFs also demonstrated that the spatial 668 variability of the satellite-derived LST fields (ASTER and) ETM+) was smaller than the one reported from the UAV LST 3 669 670 fields during day and night (see instants in Tables I, II and III)24 671 The spatial resolution of the different sources of LST might \$25 the main reason for such differences in the spatial variability 6 672 of the LST fields, especially in a heterogeneous area as the7 673 674 UIB Campus studied here. Nevertheless, the spatial variability 8 ties of the satellite LST fields for any of the studied IOP were similar to the one reported by the UAV-TIR camera for the overcast day (IOP 4, Figure 8b) demonstrated that the spatial resolution of these satellite images was not enough to properly describe the heterogeneity of the LST in the UIB Campus.

# IV. EFFECTS OF THE ADVECTION IN THE SURFACE ENERGY BUDGET AT DIFFERENT SCALES

The surface energy budget is intended to take into account all the energy exchanges taking place in a volume across the atmosphere-surface interface. As indicated in [25], it can be formally derived from the equation of the evolution of the air temperature. In this framework, the four main traditional terms (Rn, H, LE and G) would only balance if no other thermal sources or sinks existed, implying no thermal advection and surface homogeneity. Therefore, once the surface heterogeneity occurs [69], thermal advection will take place, and the usual 4-term energy budget will not close anymore. The SEB was reformulated in [25] as:

$$R_n + H + LE + G + Imb = 0$$
 (5)

where the imbalance (Imb) is the sum of the heat storage in the volume, the biological thermal exchanges, the temperature tendency, the thermal advection and any other unaccounted factor (Imb=S+B+TT+A+Ot). The authors found averaged values of the Imb, for a 2-year series in a locally flat heterogeneous terrain, between 10 and 30% depending on the time of the day and of the year, in good accordance with the estimations of [26]. Then, [19] found that the order of magnitude of the advection term was comparable to the imbalance, provided that persistent hectometer-scale heterogeneities were present around the SEB site.

The present study provides an opportunity to estimate the advection term using a homogeneous data source (the TIR on the UAV) and to check how this estimate changes with the scale of the heterogeneities. The use of the LST derived from the UAV to estimate the different terms in the SEB has been recently explored [21] and results showed variability in the turbulent heat fluxes in a heterogeneous region (100 m) but assumed the imbalance as zero.

Figure 10 shows the daily evolution (21/07/2016, IOP5) of the 4 main fluxes plus the imbalance term measured every 30 minutes (5) in the SEB station installed in the UIB Campus (yellow circle in Figure 1). The Imb shows percentages with respect to the Rn+G combined fluxes of approximately 10-20 % between 0000-0600 UTC and between 1800-2400 UTC, which were greater than the percentage values of H and LE. Between 0600-1800, the Imb represents 20-30 % of the combined fluxes, just exceeded by the H term. Thus, the Imb is a very important term to take into account in the studies of energy fluxes exchange, and the causes of this term deserve to be analyzed.

The importance of the advection term (A) in the SEB equation, which represents the effect of the motions of timescales longer than the turbulence-averaged ones, was addressed in [19]. This advection term is expressed as:

729 
$$A = \rho C_p \Delta z U \frac{\Delta T}{\Delta x} \tag{6784}$$
785

where  $\rho$  is the air density, Cp is the heat capacity at constant  $\frac{786}{nt_{-}}$ 730 pressure, u is the wind speed,  $\Delta T$  is the thermal gradient  $be_{-}^{287}$ 731 tween two points and  $\Delta x$  the distance between them,  $\Delta z$  is the 732 733 measurement height (2 m) of the atmospheric variables. The θŌ 734 expression makes an arbitrary choice taking the height of the measurement as 2 m, to be consistent with the box-approach  $\frac{761}{100}$ 735 made in [25]. Therefore, the values are just indicative, where  $\frac{792}{793}$ 736 737 as the changes are more meaningful.

794 738 The study [19] analyzed the order of magnitude of the 739 term, obtained with data from different sources during the 740 BLLAST campaign [61]. Several simplifications were base 741 on a strong hypothesis, such as neglecting the vertical advect 98 742 tion or taking the wind speed of 1 m/s as the order of magni tude in the surface layer. It was found that the contribution of  $\frac{799}{2}$ 743 Ō0 744 the scales greater than a kilometer to the advection term Q1 745 very small, whereas at the meter scale, it is too large to <u>0</u>2 746 meaningful, and its effects are probably already included Q3 747 the turbulent sensible heat flux. As long as the scales represent Õ4 748 persistent motions (therefore non-turbulent), the scales Ö5 749 between a few decameters or some hectometers, may explain 8Ö6 750 significant part of the imbalance.

In this work, we discuss if the imbalance measured at  $\frac{807}{100}$ 751 SEB station at the UIB Campus for the IOP 5 (21/07/2016) to 752 the advection terms calculated by (6) with the LST product of 753 the UAV-TIR camera at different spatial resolution, after degrading the original resolution of 2 m x 2 m to decameter of  $^{810}$ 754 755 756 hectometer resolutions. In comparison with [19], in this study, 1 the LST variability is taken approximately 4 times larger 81/2 757 758 the daytime than the temperature at 2 m and is similar in the3 759 nighttime, according to (Gemma Simó personal communical4 760 tion). Therefore, the daytime values are overestimated. The 5  $\rho C_p \Delta zu$  term in (6) is calculated with the actual wind speed 6 761 762 data, whereas the absolute values of the temperature gradients7 763 are computed between each pixel and the four (north, south,8 east and west) contiguous pixels, taking the mean value 819 764 those four gradients. As a supplement, the advection terms 20765 766 retrieved with the LST product of the Landsat 7-ETM+ and 1 767 ASTER (daytime and nighttime overpass) are also calculated22 The signs of the wind speed and of the temperature gradiena3 768 are ignored, and the corresponding values are upper value4 769 825 770 estimations.

Figure 11 shows the imbalance measured between 0400 and 6 771 772 2200 UTC of the IOP5 (21/07/2016) in the SEB station. 847 shows values below 25 W/m<sup>2</sup> at night with an absolute min 28 773 mum near dusk. As the morning progresses, the imbalan&29 774 775 increases to values near 100 W/m<sup>2</sup> during the central part  $\delta t^{0}$ 776 the day, decreasing gradually during the afternoon. This ev 831 777 lution is very well mimicked by the estimated advection terena 778 computed at a resolution of 200 m x 200 m, which captures<sup>3</sup> 779 well the dusk minimum, the morning increase and the after34 780 noon decrease. The value of advection increases with decreas<sup>35</sup> 781 ing resolution until 200 m x 200 m, leveling off for lower6 resolutions. Instead, in the morning, the advection term at 2807782 783 m x 200 m is uncorrelated with the imbalance, indicating that8 839

other factors could be important.

The advection terms at the decameter scale (represented by the curve at a resolution of 50 m x 50 m) increase in the afternoon compared to the morning and early afternoon. The curve at 2 m x 2 m is very similar to that at 50 m x 50 m; however, it takes much larger values, indicating that the thermal homogeneities at these scales may be handled by turbulence mixing. The values computed from ASTER, which has a resolution of 90 m x 90 m, are similar to the hectometer ones, whereas those from Landsat 7 ETM+ (at 30 m x 30 m) are larger, similar to what occurs with the corresponding degraded resolution of the UAV-TIR Camera product.

Consequently, we see that the estimated value of the advection term increases with the resolution at which it is computed and that the hectometer scales are the ones that behave qualitatively and quantitatively more similar to the imbalance of the surface energy budget, at least for a cloudless day. For an overcast case, the advection estimate would be smaller due to the low LST heterogeneities observed during the day (IOP4, Figure 9). The underestimation in the morning may be related to other intervening factors, for instance, an underestimation of the latent heat flux, confirming the conclusions reached in [19], and making more realistic advection calculations than those made in that work. However, the advection term cannot close the SEB equation (5) since it is obvious that there exist other terms that should also be taken into consideration.

#### V. CONCLUSIONS

A heterogeneous area containing different types of surfaces located in the University of Balearic Islands in Mallorca (Spain) was the target for an analysis of temperature gradients at high spatial resolutions since recent review studies have reported that land surface temperature heterogeneities within decameter and hectometer scales are not represented in the effective temperature of a kilometric heterogeneous area. New high-resolution orbiting TIR sensors are planned to be launched in the near future, and more studies need to be conducted prior to establishing application methods with such new LST products. On the other hand, a potentially optimum technology, not sufficiently exploited, is the use of an unmanned aerial vehicle carrying a TIR camera onboard. The present study combines both objectives together with the use of the currently orbiting TIR sensors ETM+ and ASTER.

The results show that such LST products, after validation, are capable of detecting significant temperature gradients in a heterogeneous area, which can reach differences, in the case of the UAV-TIR camera system, of up to 18 °C during the morning and 14 °C at night. These differences remain significant with the high-resolution satellite TIR sensors but were not seen with the medium-resolution LST product of the MODIS sensor. An application study of the LST heterogeneity effect on the horizontal imbalance registered in a surface energy balance station was also carried out by means of the estimation of an advection term that takes into account the turbulent heat fluxes produced due to temperature differences between a specific point and the surrounding pixels. The results confirmed that the LST gradients within hectometer resolutions

could partly explain the imbalance measured through the  $a^{829}_{900}$ vection fluxes; however, the finer scales (decametric or low 900resolutions) are too high, indicating that these scales are monop

- 843 likely handled by the turbulent mixing.
- From this study, it can be concluded that for heterogeneous  $\frac{904}{905}$ 844 zones, LST products from sensors of Landsat series or ASTER66 845 are suitable to stipulate the advection term, reducing if it 907 846 taken into account, but not fully explaining, the imbalance 847 produced in the SEB budget. This study also opens the applito 848 cation possibilities for decametric LST products of TIR sen11 849 sors, such as the near-future missions HyspIRI  $[62]_{22}^{12}$ 850 MISTIGRI [63] and THIRSTY [64]. However, further refine 14 851 852 ments need to be made to retrieve the actual value of the a@15 853 vection term and its influence in the imbalance, from both the  $\frac{1}{2}$ in situ and the satellite measurements. However, refinements  $\frac{1}{8}$ 854 855 in this matter need to be made in future works. 919
- The results from the present study are applicable to other pr
- Finally, the potential advantages of using a UAV in the 7
  study are worth noting since a better understanding of the L\$28
  heterogeneities produced in our area, as well as the possibilities of retrieving an LST map under cloudy conditions, was plays
  sible only with the UAV flights.
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# TABLE LIST

 Table I. Date and UTC hour of the daytime and nighttime ASTER sensor passes over the UIB Campus during the period April and November 2016.

	Date	UTC_Hour
es	9-Apr-16	1042
scen	19-Juny-16	1049
TER	5-Jul-16	1049
g AS	21-Jul-16	1049
ornin	7-Sep-16	1049
WG	10-Nov-16	1049
sər	19-Juny-16	2154
scer	5-Jul-16	2154
STER	21-Jul-16	2154
ht A	22-Aug-16	2153
Nig	7-Sep-16	2153

 Table II. Date, path/row and UTC hour of the daytime ETM+ sensor passes over the UIB Campus during 2016. Only clear-sky days are considered.

Date	Path/Row	Hour_UTC
20-Jan-16	196/33	1027
28-Feb-16	197/32	1033
9-Apr-16	196/33	1027
16-Apr-16	197/32	1033
2-May-16	197/32	1033
11-May-16	196/33	1027
18-May-16	197/32	1033
19-Jun-16	197/32	1033
28-Jun-16	196/33	1027
5-Jul-16	197/32	1033
21-Jul-16	197/32	1033
30-Jul-16	196/33	1027
15-Aug-16	196/33	1027
16-Sep-16	196/33	1027
23-Sep-16	197/32	1033
9-Oct-16	197/32	1033
3-Nov-16	196/33	1027
10-Nov-16	197/32	1033
28-Dec-16	197/32	1033

Table III	. UTC hour	of the UA	V flights	for the :	5 IOPs	carried	out	between.	June and	l July	/ 2016 a	at the	UIB	Cam	pus.

IOP number	Date	UTC Hour										
1	19-Jun-16	0400	0600	0800		1200	1400	1600	1800	2000		
2	28-Jun-16				1000	1200	1400	1600	1800	2000	2200	0230
3	5-Jul-16	0400	0600	0800	1000	1200	1400	1600	1800	2000		
4	14-Jul-16	0400	0600	0800	1000	1200	1400	1600	1800	2000		
5	21-Jul-16	0400	0600	0800	1000	1200	1500		1800	2000	2200	

# 

**Figure 1.** Location of the UIB Campus in Mallorca Island (left) and details of the study area (right), where the different types of surfaces can be observed. The yellow dot indicates the SEB station, while the red cross shows the position where the UAV was launched for all flights. Source: Google Earth.



**Figure 2.** RGB composite of the ETM+ passes over the Mallorca Island in two different orbits: a) path/row 196/33 scene on June 28<sup>th</sup> of 2016, b) path/row 197/32, July 21<sup>st</sup> of 2016.

# FIGURES LIST



**Figure 3. a)** Picture of the UAV-TIR camera ensemble prior to start the flight. b) Spectral response of the FLIR LEPTON camera in the 6-15µm range provided by [52].



**Figure 4.** Comparison of the LST measurements of the IR120 field radiometer with LST products retrieved from ASTER (Table 1) and ETM+ (Table 2) orbiting sensors and aerotransported TIR camera (Table 3). Linear trendline for the three sensors and  $R^2$  are included.



**Figure 5.** LST maps composed from the scenes registered by the TIR camera assembled in the UAV for the IOP 5 (21/07/2016) for: a) 04:00, b) 22:00, c) 10:00 and d) 12:00 UTC hours.



**Figure 6.** LST products of the simultaneous overpassing sensors on 21/07/2016 (IOP 5) between 1033-1049 UTC for: a) MODIS, b) ASTER, c) ETM+, and d) UAV-TIR camera flight. The SEB station position is shown as a green circle, and the area of pixels used for calculating the spatial gradient for Equation 6 at a 50-m resolution (pixel included in black squares) and 200-m resolution (pixel included in white squares).



**Figure 7.** LST products of the simultaneous overpassing sensors on 21/07/2016 (IOP 5) at 2153 UTC for: a) ASTER, b) ETM+, and c) UAV-TIR Camera Flight. SEB station position is shown as the green circle, and the area of pixels used for calculating the spatial gradient for Equation 6 at a 50-m resolution (pixel included in black squares) and 200-m resolution (pixel included in white squares).



**Figure 8.** Probability density functions (PDFs) computed from the LST fields over the 1-km square UIB-Campus derived from the multicopter (in lines) and from satellite images (ASTER and ETM+, in points) for several instants: (a) 0400 UTC, (b) 1000 UTC, (c) 1200 UTC, and (d) 2200 UTC. All IOPs are included in the plots. To make the PDFs for all IOPs comparable, the LSTs values in the x-axis are normalized with the corresponding mean value.



**Figure 9.** Temporal evolution of the statistical parameters computed from the probability density functions (PDFs) as those in Figure 8 but considering all IOPs: (a) standard deviation and (b) skewness. Those computed from the UAV are shown with circles, whereas the squares and triangles are those computed from ETM+ and ASTER, respectively.



**Figure 10.** Daily evolution of the four main terms in the SEB equation (Rn, H, G and LE) measured by the SEB station (yellow dot in Figure 1) and the corresponding imbalance (Imb) during 21/07/2016 (IOP 5).



**Figure 11.** Comparison along different hours of the IOP 5 (21/07/2016) of the imbalance measured in the SEB station with the order of magnitude of the advection term calculated in the same SEB station location with data of the LST map of UAV-TIR camera at the original spatial resolution (2 m x 2 m) and at two different spatial resolutions: a decametric pixel of 50 m x 50 m and a hectometer resolution of 200 m x 200 m, after degrading the original LST map.